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1 **Towards Completing the Rain Drop Size Spectrum:**
2 **Case Studies Involving 2D-Video Disdrometer, Droplet Spectrometer, and**
3 **Polarimetric Radar Measurements**

4
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ABSTRACT

Analysis of drop size distributions (DSD) measured by collocated Meteorological Particle Spectrometer (MPS) and a 3rd generation, low-profile, 2D-video disdrometer (2DVD) are presented. Two events from two different regions (Greeley, Colorado, and Huntsville, Alabama) are analyzed. Whilst the MPS, with its 50 μm resolution, enabled measurements of small drops, typically for drop diameters below about 1.1 mm, the 2DVD provided accurate measurements for drop diameters above 0.7 mm. Drop concentrations in the 0.7 to 1.1 mm overlap region were found to be in excellent agreement between the two instruments. Examination of the combined spectra clearly reveals a drizzle mode and a precipitation mode. The combined spectra were analyzed in terms of the DSD parameters, namely the normalized intercept parameter, N_w , the mass weighted mean diameter, D_m , and the standard deviation of mass spectrum, σ_M . The inclusion of small drops significantly affected the N_w and the ratio σ_M/D_m towards higher values relative to using the 2DVD-based spectra alone. For each of the two events, polarimetric radar data were used to characterize the variation of radar measured reflectivity (Z_h) and differential reflectivity (Z_{dr}) with D_m from the combined spectra. In the Greeley event, this variation at S-band was well-captured for small values of D_m (< 0.5 mm) where measured Z_{dr} tended to 0 dB but Z_h showed a noticeable decrease with decreasing D_m . For the Huntsville event, an overpass of the Global Precipitation Measurement mission Core satellite enabled comparison of satellite-based dual-frequency radar retrievals of D_m with ground based DSD measurements. Small differences were found between the satellite-based radar retrievals and disdrometers.

48 **1. Introduction**

49

50 Knowledge of the drop size distribution (DSD) at different scales and in different rainfall types
51 and rain intensities is of obvious importance in both practical radar applications as well as in
52 numerical parameterizations of the fundamental microphysical processes such as collision-
53 coalescence, drop break-up and evaporation. Due to the large variability of the DSD (Bringi et
54 al., 2003, Ulbrich 1983), it has been conventional (depending on application) to consider
55 moments of the DSD such as mass-weighted mean diameter (D_m), normalized intercept
56 parameter (N_w) and width of the mass spectrum (σ_M) as well as the shape of the normalized and
57 scaled distribution (e.g., Ulbrich and Atlas 1998; Testud et al 2001; Haddad et al. 1996; Semper-
58 Torres et al, 1994; Lee et al. 2004). Whilst the higher order moments (≥ 3) involved in
59 calculating D_m , N_w or σ_M are generally considered to be much less sensitive to the small and tiny
60 drop end of the DSD (typically diameters < 0.7 mm), both the total number concentration (zeroth
61 moment) and the shape of the distribution can be significantly controlled by the small drop end
62 which is difficult to measure accurately.

63

64 The DSD is generally measured at the surface using optical or impact-type disdrometers
65 typically averaged over several minutes to capture the distribution of the aforementioned DSD
66 parameters with rain rate. It is also well-known that most, if not all, disdrometers tend to
67 underestimate the concentration of small and tiny drops ($D < 0.7$ mm or so) because of
68 sensitivity issues and poor resolution, and – depending on the design – other instrumental factors
69 may also play a role (Tokay et al. 2001; Miriofsky et al. 2004). The accurate measurement of
70 tiny drops is important for the calculation of the total concentration of drops (N_{tot}), as well as in

71 the numerical modeling of collision-coalescence processes of rain formation and DSD evolution
72 (e.g., Meyers et al. 1997; Milbrandt and Yao 2005). For example, the probability that a large
73 drop will undergo collisions with a tiny drop is proportional to (among other factors) the
74 concentrations of the latter. Ideally, such concentrations should be measured with very high
75 resolution instruments developed for airborne applications (e.g., 2D-C or cloud imaging probe)
76 but these have been rarely used as surface disdrometers (Montero-Martinez et al. 2009).

77
78 More importantly for polarimetric radar applications, collisions between moderate-to-large drops
79 ($D > 2$ mm or so) with tiny drops ($D \sim 0.5$ mm) has long been postulated as a viable mechanism
80 of producing large amplitude oscillations in the larger drop (possible precursor to drop break-up)
81 that can be sustained against viscous dissipation (Beard et al. 1983). The excess kinetic energy
82 due to collisions (or simply collision kinetic energy) which can force such oscillations is
83 proportional to (among other factors) the volume of the tiny drops, so their sizes should also be
84 measured with high resolution. Johnson and Beard (1984) determined that the most energetic
85 collisions were those between moderate-to-large drops ($D > 2$ mm) and tiny drops in the range
86 ($D \sim 0.3$ – 0.8 mm). This re-emphasizes the importance of measuring the tiny drops with higher
87 resolution than is possible with current surface disdrometers.

88
89 The gamma DSD model (Ulbrich, 1983) is widely used in polarimetric and dual-wavelength
90 radar applications but the shape parameter (μ , as defined in Ulbrich and Atlas, 1998.) and its
91 dependence on rain microphysics is not well-established via surface distrometer measurements
92 principally due to difficulty in measuring the concentrations at the small drop end which plays a
93 strong role in estimating the μ -parameter. Assumptions of constant μ (≈ 3), or empirical μ - Λ

94 (where $\Lambda D_m=4+\mu$) fits, or statistical methods based on fits to σ_M-D_m variations are susceptible to
95 errors which are not easily quantified (e.g., Kozu et al., 2009; Zhang et al.2003; Williams et al,
96 2014). On the other hand Testud et al. (2001) found remarkable stability of shape of the
97 normalized and scaled DSD (non-gamma model) using aircraft-based imaging probes in oceanic
98 rainfall. It is not clear if a single gamma model can be used to describe the shape of the entire
99 DSD (e.g., Able and Boutle 2012). While there is vast literature on DSD measurements based on
100 surface disdrometers or aircraft imaging probes, very few studies accurately characterize the full
101 size spectrum which needs at least 2 instruments and an overlapping size range to ensure that
102 instrumental errors are low (i.e., to ensure consistency and continuity of concentration
103 measurements in the overlap size range), and that the resulting data can be used for physical
104 interpretation of the DSD shape and variability.

105
106 In this paper we describe DSD data collected with ‘side-by-side’ collocation of the
107 Meteorological Particle Spectrometer (MPS; Baumgardner et al. 2002) with a 3rd generation,
108 low-profile, 2D-video disdrometer (2DVD; Schönhuber, et al., 2008) to enable us to characterize
109 the concentration of the tiny drops with very high resolution (50 μm) with the MPS, and the
110 same for larger drops (with resolution of 170 μm) from the 2DVD. Our objective then is to
111 combine the MPS and 2DVD data to form a composite DSD with high resolution at the small
112 drop end provided by the MPS and good resolution provided by the 2DVD for moderate-to-large
113 drops. So far, measurements at two locations have been carried out, namely Greeley, Colorado,
114 and Huntsville, Alabama, and we report here on observations and analysis from two long
115 duration events from the two sites.

2. Instrumentation and experimental set-up

a. The Two Campaigns

The Greeley campaign took place from April to October 2015 and the Huntsville campaign started in March 2016. The same MPS and the 2DVD instruments were used in the Greeley campaign and the Huntsville campaign.

At the Greeley site, both instruments were conveniently installed within a 2/3-scaled double fence inter comparison reference (small DFIR, the standard adopted by the National Weather Service for snow gages) windshield, located at about 13 km SSE from the CSU-CHILL S and X-band polarimetric radar site (Bringi et al., 2011). The sensor areas of the disdrometers were set at a height which is 13 inches below the top of the inner fence. The small DFIR had been originally built for a snow observation campaign and had proven to be effective in substantially reducing wind speeds (Fig. 17 in Notaros et al., 2016). A Pluvio weighing bucket-type rain gauge was also installed within the wind fence. This was a 2nd generation weighing type rain gauge manufactured by OTT with a 200 cm² collection area that utilizes a highly precise load-cell to enable study of rainfall amounts as little as 0.1 mm with an accuracy of 0.2% (OTT Messtechnik GmbH 2010).

In Huntsville, Alabama a similar small DFIR located at the National Space Science and Technology Center (NSSTC) on the campus of the University of Alabama in Huntsville (UAH) housed the MPS and the 2DVD. The sensor areas were also set at the same height as in the

139 Greeley campaign. The site is located 15 km from the UAH/WHNT-TV Advanced Radar for
140 Meteorological and Operational Research; (ARMOR), which is a C-band polarimetric radar
141 (Petersen et al. 2005, <http://www.nsstc.uah.edu/armor/>). Fig. 1(b) shows the ground instruments
142 and the small DFIR configuration at the Huntsville site.

143

144 *b. Small Drop Measurements with MPS and Overlap with 2DVD*

145

146 The MPS uses a linear array of 64 photodiodes to measure the shadow images of particles falling
147 through a collimated laser beam. The concepts of the technique were originally introduced by
148 Knollenberg (1970), and later by Baumgardner et al. (2002). This instrument has 50 micron
149 resolution and is suitable for measuring small drops. The size range is 50 microns to 3.1 mm and
150 its sampling area is 6.2 cm². The 2DVD on the other hand has a much larger 10 by 10 cm² sensor
151 area (Schönhuber, et al., 2008) but the pixel resolution for the front and side view (silhouettes) is
152 around 170 microns.

153

154 The 2DVD is a well-established disdrometer that uses two optical cameras to measure the size,
155 shape and fall velocity of individual raindrops (Schönhuber, et al., 2008). Of all the
156 disdrometers, this instrument has been established as the most suitable instrument for measuring
157 the large drop end of the DSD spectrum (Gatlin et al., 2015). On the other hand, this instrument
158 does not reliably measure the drop concentration for drop diameters less than about 0.6 mm; in
159 fact, it tends to underestimate N(D) for these small drops (Tokay et al., 2001). The problem is
160 related to lowered sensitivity to small and tiny drops, the associated difficulty in matching of
161 these drops from the two camera images and to finite instrument resolution.

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The MPS is a high resolution instrument for drop imaging and measurement of the DSD specifically designed for fixed site operation (see Fig. 2 and Table 1). It was developed in early 2000 to measure drizzle for the National Weather Service. Fall speeds are measured with the MPS after sizing the horizontal dimension (or the width ‘ a ’ in Fig. 2b) and dividing by the time taken to traverse the photo-detector array (spherical shape is assumed, i.e., vertical dimension is equal to ‘ a ’). Table 1 gives some of the important technical specifications of the MPS and the 2DVD.

The fall speed accuracy of the MPS depends primarily on the digitization error ($\pm 25 \mu\text{m}$), and according to the manufacturer it is 10% for $D = 0.25 \text{ mm}$ and 1% for $D = 1 \text{ mm}$. The factors that determine the accuracy of the 2DVD for size, fall speed, and axis ratio are given in Schönhuber et al. (2008), Kruger and Krajewski (2002), and Thurai and Bringi (2005).

The effective measurement area of the MPS decreases with increasing drop width (“entire in” images; Heymsfield and Parrish 1978) and is a factor of ≈ 30 smaller relative to the 2DVD for a measured drop width of 1.5 mm. This increases the sampling error for estimation of the concentration of drops with $D \approx 1.5 \text{ mm}$ by a factor of $\sqrt{30} \approx 5.5$. In our application, we will utilize the MPS for measurement of small drops with $D < 1.2 \text{ mm}$ and to compare the measurements with the 2DVD in the overlap region of $D \approx 0.7\text{--}1.2 \text{ mm}$ to ensure consistency of observations. The method of deriving the drop size distribution from the MPS is summarized in the Appendix.

3. The Greeley Campaign

The event considered in this paper occurred on 17 April 2015, soon after the MPS installation at the small DFIR site at Greeley. This event was part of a mid-latitude synoptic scale cyclone that had produced fine drizzle, light precipitation, cold rain, rain bands (both stratiform and convective in nature) as well as thunderstorms towards the end of the event (Thurai et al., 2015). The CSU-CHILL S-band radar scans were made at regular and closely spaced time intervals and consisted of surveillance plan position indicator (PPI), sector PPI, and range-height indicator (RHI) scans. The preprogrammed scan sequence included (a) a 360 scan at 10 deg elevation, (b) two-sweep RHI scans over the disdrometer site, and (c) a one sweep (1.5 deg elevation) PPI sector volume centered over the disdrometer site, which were repeated every 5 min and 27 seconds.

(a) *Ground instrument data*

(I) Fall velocities

The 17 April 2015 event (Greeley Event) was an intermittent but long duration event which produced a variety of rain types over a period of 20 hours. Whilst the MPS enabled drop concentration measurements down to 0.1 mm (i.e. with at least 2 pixels), the 2DVD recorded drops as large as 5 mm associated with the (non-hail producing) thunderstorm. Fall velocities showed a clear trend with drop diameter, in agreement with the expected Gunn-Kinzer variation, but with an adjustment factor appropriate for the 1.4 km altitude for Greeley. Fig. 3(a) shows the comparisons for all drops. Fig. 3(b) shows the distribution of velocity for drops with equivalent

208 spherical diameters (D_{eq} or D) of 2.5 ± 0.1 mm. The increased fall velocities can be clearly
209 attributed to the reduced pressure at the 1.4 km height above mean sea level (AMSL).

210

211 (II) Rain rates and accumulations

212 The processed Pluvio data are shown in Fig. 4: (a) shows the 1-minute rain fall rate, (b) shows
213 the rainfall accumulation, and (c) shows the corresponding 2-hour total accumulation. The 1-
214 minute rain rate was as high as 19 mm/h (towards the end of the event), the total event
215 accumulation over the 20 hour duration was 17 mm, and the 2-hour accumulations varied
216 significantly throughout the event, ranging from 0.05 mm to 4.91 mm.

217

218 Based on the rainfall rates and other ground instrumentation data as well as the corresponding
219 CHILL scans for the entire event duration, a broad rain-type classification for each of the 2-hour
220 period (corresponding to Fig. 4(c)) was made, as given in Table 2. Note the highest rainfall rate
221 occurred during the thunderstorm period (18 - 20 UTC), and the lowest rain accumulation
222 occurred during the fine drizzle period (10 - 12 UTC). (Here we have adopted the AMS glossary
223 description of drizzle as ‘form of precipitation consisting of water droplets less than 0.5 mm in
224 diameter and larger than 100 nm.)

225

226 (III) Drop size distributions

227

228 For drop size distribution comparisons between the 2DVD-based and the MPS-based
229 measurements, we first split the entire time series event into the same 2-hour time intervals
230 mentioned earlier, starting with 02:00 UTC and ending with 20:00 UTC. Fig. 5 shows these 2-

231 hour DSD comparisons, except for the last two hours. The diamonds represent the MPS-based
232 DSD's and the crosses represent the 2DVD-based DSD's. Close overlap is seen in the 0.7 - 1.2
233 mm drop diameter range. For over 95% of the cases, the fractional differences between the MPS
234 and the 2DVD drop concentrations (on a log scale) in this diameter range was less than 10%, and
235 moreover the overall average was found to be -3.8% which is very close to zero, indicating that
236 there is no systematic bias. The log-log scale was used to focus on the small drop size range. The
237 concentration of smaller drops was underestimated by the 2DVD relative to the MPS, as
238 expected. However, the 2DVD measurements of moderate to large drop sizes (drop volume is
239 based on two orthogonal views) can be considered to be more accurate than the MPS, since the
240 latter assumes *apriori* spherical drop shapes.

241
242 As a result of the consistency between the two instruments demonstrated in Fig. 5, the full DSD
243 spectra were constructed based on the drop concentrations from the MPS for $D_{eq} < 0.7$ mm and
244 the 2DVD-based drop concentrations for $D_{eq} \geq 0.7$ mm. Examination of Fig. 5 reveals two
245 different modes, (i) a drizzle component for $D_{eq} \leq 0.5$ mm and (ii) a precipitation mode for
246 larger diameters (starting near or about the shoulder region especially noticeable in the 04:00-
247 06:00 UTC panel). Such modes have been previously identified from aircraft imaging probe (2D-
248 Cloud and 2D-Precipitation) data collected in warm rain clouds analyzed by Able and Boutle
249 (2012). In fact, their combined spectra from the 2D-C (similar to MPS) and 2D-P (similar to
250 2DVD) are very similar in shape to Fig. 5. They also show that an exponential shape forms a
251 good fit to the precipitation mode portion of the combined spectra (easy to see as a straight line
252 in a more conventional semi-log plot of the DSD). Our MPS-2DVD results are consistent with
253 their analysis in spite of different instruments, time integration and meteorological conditions
254 (in-cloud oceanic warm rain versus continental spring-time surface precipitation).

255
256 Fig. 6 shows the comparisons of two DSD parameters based on the 1-minute DSDs from the
257 combined spectra (shown as black dashed line) and those solely from 2DVD (grey crosses). The
258 two parameters are the mass-weighted mean diameter (D_m) shown in panel (a) and the standard
259 deviation of the mass spectrum (σ_M) shown in panel (b), as defined in Ulbrich and Atlas, (1998).
260 During the thunderstorm period (18:00 – 20:00 UTC) rapid fluctuations can be seen in both
261 parameters, and this correlates well with the rapid fluctuations in rain rates from Pluvio
262 measurements shown earlier in Fig. 4(a).

263
264 The inclusion of the small drops from the MPS in the combined spectra results in a decrease in
265 D_m and an increase in σ_M . The resulting variation of σ_M versus D_m for the combined spectra
266 (diamonds) is shown in panel (c) and compared with those based on 2DVD data alone (crosses).
267 The dashed line in panel (c) represents the best-fitted power law equation - using log-linear
268 model - for the variation based on the 2DVD data alone. Note the power-law fitted equation is
269 close to that given in Thurai et al. (2014) and Williams et al. (2014) who used 2DVD data alone
270 from a long measurement campaign in Huntsville, AL. For the combined DSDs, it was not
271 possible to fit a representative power-law equation (of the form $\sigma_M = \alpha D_m^\beta$) for the entire dataset
272 primarily because the σ_M estimates become noisy for low D_m , e.g., between 10 – 12 UTC in
273 panels (a) and (b). If the fit is performed for the data with $D_m \geq 0.5$ mm, the fitted equation
274 becomes $\sigma_M = 0.48 D_m^{0.94}$ for the combined spectra which is significantly different from the
275 fitted equation using the 2DVD spectra alone ($\alpha=0.29$, $\beta=1.44$).

276

277 Fig. 7 shows a set of D_m histograms based on the 2DVD-based DSDs and the combined DSDs.
278 The three top panels correspond to the 2-hourly periods of 02:00–04:00, 08:00–10:00 and 18:00–
279 20:00 UTC, corresponding to convective rain, light rain or stratiform rain, and thunderstorm. The
280 following summarizes some pertinent points:

281

- 282 i. Histograms from the combined spectra show lower values of D_m than those from
283 the 2DVD alone.
- 284 ii. Light stratiform rain produces histograms with lower D_m than convective rain (as
285 expected).
- 286 iii. There is considerable difference between the modal values of D_m in light rain
287 (mostly non-overlapping histograms)
- 288 iv. The thunderstorm period histograms are similar for larger D_m but the MPS-2DVD
289 based DSDs have more cases with low D_m (≤ 0.5 mm).

290

291 The lower panels in Fig. 7 show the D_m histograms classified in terms of four rain rate intervals.
292 The very low rain rates with $R < 0.5$ mm/h (including drizzle) shows skewed histograms for both
293 cases, but the combined DSDs give rise to noticeably lower D_m values. The histogram shows a
294 peak of around 0.15 mm for the combined DSDs versus 0.6 mm for the 2DVD-based DSD's.
295 The histograms become more similar for the higher rain rates, exhibiting peaks at around 1 mm
296 for the $1 < R < 5$ mm/h interval. For $R > 5$ mm/h, the peaks are around 1.15 mm, but the total
297 number of points was only 27.

298

299 Fig. 8(a) compares the D_m calculated using only the 2DVD spectra versus those using the
300 combined spectra. Compared with the [1:1] dashed line, the bias is evident, and in almost all
301 cases, the 2DVD-only spectra tend to overestimate D_m , which is to be expected, but the
302 overestimation is higher when $D_m < 1$ mm (i.e., for DSDs where small drops play a more
303 dominant role). Note, also, that the D_m calculated from the 2DVD-only spectra shows a floor at
304 0.5 mm since the small drop concentrations are strongly underestimated.

305
306 In terms of rain accumulation, the addition of the small drops from the MPS provided small but
307 significant improvement in the agreement with the collocated Pluvio data for the two convective rain
308 event periods (02-04 UTC, and 04-06 UTC). Table 3 shows the comparisons for the 2-hour period. Also
309 shown are the comparisons for the 18-20 UTC time period which included modest thunderstorm activity.
310 In all three cases, the 2-hour accumulations from the composite MPS-2DVD DSDs show better
311 agreement with Pluvio data. For other two hour periods, accumulations were less than 2 mm.

312
313 *b) S-band CHILL Radar observations*

314
315 As mentioned earlier, the CHILL S-band radar scans were made over the instrumented site at
316 regular and closely spaced time intervals (< few minutes). From the surveillance and sector PPI
317 scans, values of Z_h and Z_{dr} over the instrument site were extracted (az: 171.5 deg, range 13 km).
318 Only the radar pixel directly above the instrument site was considered, and no spatial averaging
319 was done. The radar pulse volume was centered at ~310 m above the disdrometer site (which
320 was around 30 m higher than the radar site). Fig. 9 (a) and (b) shows the variation of these values
321 versus D_m derived from the 1-minute DSDs (but smoothed over 3-minutes) with the S-band Z_h

322 and Z_{dr} extracted over the instrument site. The grey crosses represent the D_m values obtained
323 from the 2DVD spectra alone and the black diamonds represent those derived from the 2DVD-
324 MPS combined spectra.

325
326 Some important points can be noted from Fig. 9(a) and (b). First, the $Z_{dr} - D_m$ variation does
327 indeed get affected by including MPS measurements of small drops, particularly for low D_m
328 values. Second, when D_m goes below 0.5 mm, the S-band Z_{dr} becomes very close to 0 dB and
329 exhibits very little sensitivity to further lowering of D_m (to be expected as the small drops are
330 close to spherical in shape). On the other hand, Z_h exhibits greater sensitivity to changes in D_m
331 even below 0.5 mm for the combined DSDs. Thus, for events with low D_m (<0.5 mm) the
332 combined spectra/CHILL radar data suggests the appropriateness of using both Z_h and Z_{dr} to
333 retrieve D_m (as opposed to using Z_{dr} alone), see Thurai et al. (2012).

334

335

336 **4. The Huntsville Campaign**

337

338 *a) Event description*

339

340 Huntsville has a very different climate to Greeley, and its altitude is 200 m AMSL compared
341 with 1.4 km AMSL for Greeley. The climate of northeastern Colorado is much drier and cooler
342 on average than that of northern Alabama. Huntsville receives an average of 138 cm of
343 precipitation each year whereas Greeley receives less than 38 cm each year. Greeley has a daily
344 mean temperature that is 4 degrees cooler than Huntsville. According to the Köppen-Trewartha

345 climate classification system (Trewartha and Horn 1980), this labels Greeley, CO as a semi-arid
346 type climate, whereas Huntsville, AL is a humid subtropical type climate (Belda et al. 2014).

347
348 The Huntsville event considered in this paper occurred on 11 April 2016, and consisted of
349 precipitation associated with the mesoscale vortex of a developing squall line that moved across
350 northern Alabama between 1700 to 2300 UTC and produced over 25 mm of rainfall in the
351 Huntsville area. This event was sampled by the MPS and 2DVD just after they had been
352 installed within the small DFIR. The ARMOR radar was performing PPI scans over these
353 disdrometers, and the Global Precipitation Measurement (GPM) mission Core satellite (Hou et
354 al. 2014) made an overpass of northern Alabama near the end of this precipitation event.

355
356 *b) Ground-based measurements*

357
358 Fall velocity measurements from the 11 April 2016 event are shown in Fig. 10(a). Once again the
359 dashed line represents the Atlas et al. (1973) fitted equation to the G-K data at sea level. The
360 2DVD measurements show much closer agreement to this variation than the Greeley data shown
361 earlier in Fig. 1(a). However, note the more intense color contours lie slightly higher than the
362 dashed line, which can be explained by the 200 m altitude above sea level. Fig. 10(b) shows the
363 histograms of vertical velocity specific for all drops with D_{eq} values of 2.5 ± 0.1 mm, whose
364 mode closely agrees with the expected fall velocity of 7.3 m s^{-1} .

365
366 For drop size distribution comparisons between the 2DVD-based and the MPS-based
367 measurements, the time series event was split into 1-hour time intervals, starting at 17:00 UTC

368 and ending at 23:00 UTC. Fig. 11 shows these hourly DSD comparisons. The black diamonds
369 represent the MPS-based DSDs and the black crosses represent the 2DVD-based DSD's. Fig. 11
370 shows similar features to Fig. 5, that is, close overlap in the 0.7 - 1.2 mm drop diameter range
371 between the MPS-based DSDs and the 2DVD-based DSDs, but again for smaller drops, the
372 2DVD underestimates the drop concentration compared with the MPS.

373
374 As with the Greeley data analysis, the combined spectra were constructed based on the drop
375 concentrations from the MPS for $D_{eq} < 0.7$ mm and the 2DVD-based drop concentrations for
376 $D_{eq} \geq 0.7$ mm. As discussed in Section 3.1(c), the two modes identified by Able and Boutle
377 (2012) are quite evident in Fig. 11—a drizzle mode for diameters < 0.5 mm, and a precipitation
378 mode starting around 0.7-1 mm (i.e., the shoulder region) and extending to the largest sizes.
379 These two modes are actually more prominent in Fig. 11 as opposed to Fig. 5 perhaps due to the
380 expected prevalence of warm rain processes in the Huntsville event, which had a 0°C level
381 around 3 km AGL, as opposed to dominance of ice phase processes in the Greeley event, which
382 had a 0°C level much lower (for example at 05:44 UTC, the LDR from a 10 degree VAD scans
383 had shown extraordinarily clear melting layer around 6 km range as in Fig. 3 in Thurai et al.,
384 2015) which gives a melting layer height of around 1 km AGL).

385
386 Fig. 12 (a) and (b) show, respectively, the time series comparisons of D_m and σ_M derived from
387 the 1-minute DSDs from the combined data from MPS and 2DVD (shown as black diamonds)
388 and those from 2DVD data alone (grey crosses), over a period of 4 hours. The same trend as the
389 Greeley results is seen, that is, the MPS-2DVD combined spectra give rise to slightly lower D_m
390 and larger σ_M relative to using the 2DVD spectra alone. Panel (c) of Fig. 12 shows the
391 corresponding effect on N_w . The higher concentration of small drops in the combined spectra

392 results in an increase in N_w . Note that the definition of the normalized intercept parameter
393 follows Testud et al. (2001) (which is independent of the gamma assumption) and, except for
394 constant terms, is proportional to the ratio of rain water content to D_m^4 . The increase in the total
395 number concentration will be even more significant (not shown here).

396

397 *c) ARMOR radar data*

398

399 The radar used for the Huntsville campaign is the C-band ARMOR radar, located 15 km from
400 the ground instrumentation site. The ARMOR scanning strategy for the 11 April 2016 event
401 consisted of plan position indicator (PPI) type (i.e., radar antenna rotates 360 degrees in azimuth)
402 scans with a repeat cycle of every 2.5 minutes. From these scans, the radar data over (and
403 surrounding) the disdrometer site were extracted (52 deg azimuth, 15 km range, and once again
404 only the radar pixel directly above the instrument site was considered, and no spatial averaging
405 was done.). The chosen elevation angle was 1.3 deg. Given that the half-power antenna
406 beamwidth is close to 1° , the cross-beam resolution will be around 250 m at the range of 15 km.
407 The height of the radar pixel above ground will be around 340 m.

408

409 The Z_h and Z_{dr} data extracted from the ARMOR PPI scans over the disdrometers are shown in
410 Fig. 13 (a) and (b), respectively, as a time series for the same 20:00 – 24:00 UTC time period. In
411 panel (b), the D_m values obtained from the combined DSDs are also included (the same as the
412 diamonds in Fig. 12a). One can see good correlation between the ARMOR Z_{dr} values and the
413 combined disdrometers-based D_m values.

414

415 The correlation between Z_{dr} and D_m is better depicted in Fig. 13c as a scatter plot which shows
416 ARMOR Z_{dr} versus the ground based D_m data from 2DVD spectra as well as the combined
417 spectra. The C-band Z_{dr} is more sensitive to D_m change than at S-band. However, the D_m values
418 for the Huntsville event did not go below 1 mm, and as noted in Thurai et al. (2012), Z_{dr} alone is
419 sufficient to estimate D_m for such cases. Note that for a given radar measured Z_{dr} , the D_m from
420 the combined spectra is typically biased low relative to the D_m from the 2DVD spectra (around a
421 few tenths of a mm at Z_{dr} of 1.5 dB). This trend is noted even in the presence of radar
422 measurement errors inherent in the scatter in panel (c).

423

424 *d) GPM overpass*

425

426 An overpass of the GPM Core satellite during this Huntsville event enabled us to examine the
427 performance of Version 4 of the DPR Level 2 algorithm (2ADPR), which assumes a fixed $\mu = 3$
428 in order to retrieve D_m and N_w from the attenuation corrected reflectivity values computed at
429 the two frequencies (Iguchi et al. 2016). The D_m value from the DPR bin closest to the
430 disdrometer site was 1.9 mm, and the average from this bin and the surrounding bins was 1.8 mm
431 with a standard deviation of 0.1 mm (Figure 14a). These values were derived from DPR
432 measurements at 23:31:44 UTC. Over a five minute period around this time, the average D_m
433 values from the 2DVD and combined MPS-2DVD was 1.73 mm and 1.61 mm, respectively (Fig
434 12a). The average N_w computed by the 2ADPR Version 4 algorithm was lower than that
435 measured by the disdrometers. The nine bin average N_w from the DPR (Fig. 14b) was 828 m-
436 3mm-1 (2.92 in log10 units) with a standard deviation of 189 m-3mm-1, whereas for the 2DVD
437 and combined MPS-2DVD the average N_w over the five minute period was 1348 (log₁₀=3.13)

438 and $1952 \text{ m}^{-3}\text{mm}^{-1}$ ($\log_{10}=3.29$), respectively. The σ_M measured from the MPS-2DVD over this
439 five minute period was 0.76 mm, which for a gamma DSD yields a μ value of 0.49.

440
441 Finally, in Fig. 14(c) we show the 10 minute DSD measurements from the MPS and 2DVD
442 during the GPM overpass time in order to illustrate the DSD agreement in the overlap region for
443 a finer time resolution (rather than over a 1 hour or two hour period). Although the MPS data are
444 somewhat more noisy, these two DSDs once again merge rather well in the 0.7 – 1.2 mm diameter
445 region.

446 447 **5. Discussion**

448 449 *(a) DSD shape*

450
451 In the past, DSD measurements at a given site have been carried out largely with the same type
452 of instrument, most of which can measure across a similar range of diameters (e.g. Parsivel
453 disdrometer, Joss-Waldvogel disdrometer and/or 2DVD, etc.). Krajewski et al. (2006) have
454 compared DSD measurements from different instruments which were located close to one
455 another and found considerable instrument-to-instrument differences which made it difficult to
456 study the natural variations in DSD at short spatial scales (< few hundred meters). Results
457 reported herein from two rain events in different climatologies show that there is very close
458 agreement between the 2DVD and the MPS spectra in the overlap region (0.7–1.5 mm drop
459 diameter) giving confidence that the combined spectra can be used to characterize the entire
460 DSD more accurately than hitherto possible with single instruments. The fact that both

461 instruments were installed within identical DFIR wind shields may have been responsible in that
462 the wind-induced effects could have been significantly reduced, especially for the MPS. On
463 average the mean wind speed inside the DFIR was reduced by a factor of three or more relative
464 to the environment outside the fence at the Greeley site (for example, Fig. 17 in Notaros et al.
465 2016). This allowed us to operate the MPS without its wind vane inside the small DFIR so the
466 laser beam was oriented parallel to the expected environmental mean wind direction at both sites
467 to help mitigate size distortion that can arise due to the horizontal motion of the drops.

468
469 The combined MPS-2DVD spectra from both the Greeley and the Huntsville sites have clearly
470 shown that (in the events analyzed herein and over 1-2 h time integration), the concentration of
471 small drops does increase significantly with decreasing drop diameter ($D < \sim 0.5$ mm) and is
472 consistent with the drizzle mode identified by Able and Boutle (2012). Our combined MPS-
473 2DVD spectral shapes are also consistent with what they identified as the precipitation mode for
474 sizes > 0.7 mm which in the log-log plots of $N(D)$ versus D , starts with a well-defined ‘shoulder’
475 region near 1 mm and curving convex downwards for larger sizes. Able and Boutle (2012) also
476 found that the exponential function provided a good fit to their data for the precipitation mode
477 (also consistent with visual inspection of our combined spectra). It is worth mentioning that their
478 data were acquired with aircraft-mounted 2D-Cloud and 2D-Particle probes in oceanic warm rain
479 cumulus clouds. The precipitation that was observed during the event in Huntsville was similar
480 in that it was largely dominated by warm rain processes and characterized by relatively weak
481 rainfall rates. Furthermore, the shape of the combined MPS-2DVD DSDs, especially during the
482 Huntsville event, resembles that of the bi-modal DSDs produced by simulations of raindrop
483 collisions (e.g., McFarquahar 2004; Straub et al. 2010). This suggests that collision-induced

484 breakups were responsible for shaping the observed shoulder region, which was more prominent
485 in the Huntsville warm rain event and during times of thunderstorms observed in the Greeley
486 campaign.

487
488 Time integration of the spectra over 1-2 h clearly brings out the systematic differences when
489 comparing 2DVD-only and the combined MPS-2DVD spectra. The combined spectra indicate
490 that a gamma model could not possibly fit the entire size range at either of the two
491 climatologically different sites. However, the gamma model with parameters (N_w , D_m , μ ;
492 Illingworth and Blackman 2002; Testud et al. 2001) has been used to describe the shape of the
493 aforementioned precipitation mode of the spectrum using data from 2DVD primarily for radar
494 applications (e.g., Bringi et al. 2003; Brandes et al. 2002; Williams et al. 2014). The fitted μ -
495 values showed a broad distribution with mean values between 3-5. To better characterize the
496 small drop end, Thurai et al (2014) describe the use of single camera data (from 2DVD; using the
497 same methodology as the MPS except for poorer resolution) to re-adjust the standard 2DVD-
498 derived concentrations for $D < 0.6$ mm. After such adjustment, the fitted μ values were found to
499 be significantly lower (μ between -2 to 2) when compared with the much larger and positive μ -
500 values for non-adjusted DSDs. One of the example cases was a light precipitation event in
501 Emasalo, Finland, where the adjusted 2DVD-based DSDs had been compared to the DSD
502 measurements made with the high resolution (25 μm) Cloud Imaging Probe on the Wyoming
503 King Air aircraft (during a spiral descent over the 2DVD). The 2DVD-adjusted concentrations
504 were found to be in good agreement with the airborne data for small drops. Both showed much
505 higher concentrations of small drops, similar to the drizzle mode found with the MPS
506 measurements both in Greeley and Huntsville, as well as those given in Abel and Boutle (2012).

507

508 Our new observations also point out that the earlier studies conclusions (e.g. Willis, 1984;
509 Ulbrich, 1985; Vivekanandan et al., 2004) regarding truncation errors for gamma model DSD's
510 may no longer hold for some integral parameters, especially those related to the lower order
511 moments and shape (or breadth) of the DSD. Furthermore, it may be that if a mathematical
512 model is required to represent the entire DSD, then a single gamma model with a triplet of
513 parameters may not be sufficient to fully represent the DSD for these properties. It may be
514 necessary to consider other models, including mixed models where one model is used for the
515 drizzle mode and another for the precipitation mode. This points to the potential need for
516 additional work on modeling DSD's across the full spectrum of measured drops sizes and to
517 assessing errors associated with those new models. Another, more attractive, formulation is the
518 generalized gamma function, as considered for example by Auf der Maur, 2001, and later by Lee
519 et al., 2004 who have illustrated a sample of possible shapes that can be represented by this
520 function. The flexibility of this method may well be suitable for describing the full DSD spectra
521 reported in this paper.

522

523 *b) Polarimetric radar retrievals*

524

525 As we saw earlier in sections 3 and 4, the higher concentration of small drops results in lower D_m
526 and higher σ_M . This has two implications. Firstly, the variation of D_m with Z_{dr} will be different
527 (i.e., for a given radar measured Z_{dr} , the D_m values are slightly lower for the combined spectra)
528 but as indicated by the Greeley results, this negative bias in D_m becomes more significant for low
529 rainfall rates. The second implication is that the σ_M versus D_m variation is significantly modified

530 when the more accurate small drop concentrations are included in the DSDs. More importantly,
531 the ratio σ_M/D_m becomes amplified due to the combined effects of increase in σ_M along with a
532 decrease in D_m for the combined spectra. Since for a gamma model $\sigma_M/D_m=(4+\mu)^{-1/2}$ it follows
533 that the “effective” μ will be significantly reduced for the combined spectra (the notion of
534 “effective” μ is introduced since the combined spectra in general would not follow the gamma
535 model, and as mentioned earlier, a better representation would be the generalized gamma
536 function). This amplification of σ_M/D_m is similar to truncating the spectra at the small drop end
537 due to instrument limitations (Ulbrich and Atlas 1998). Our combined spectra results suggest that
538 polarimetric or dual frequency retrieval algorithms that assume a constant μ value (typically $\mu\approx 3$)
539 or a μ - Λ relation (Λ is the slope factor in the gamma model, e.g., $\Lambda D_m=4+\mu$) or use the ratio
540 σ_M/D_m to estimate μ statistically (e.g., Kozu et al., 2009; Zhang et al., 2003; Williams et al.,
541 2014), may need further evaluation. Note these are some of the assumptions which are used for
542 the DSD retrievals from the GPM Dual-frequency Precipitation Radar (DPR; Hou et al., 2014,
543 Munchak and Tokay, 2008). Finally, whereas it is self-evident that the total number
544 concentration will be much higher for the combined spectra relative to the 2DVD-only case, the
545 normalized intercept parameter N_w being proportional to W/D_m^4 will also be amplified due to the
546 D_m being raised to the 4th power.

547

548 In the past, the estimation of D_m from S-band polarimetric radar has only used Z_{dr} . Our results
549 from the Greeley campaign show that for low rainfall rates the S-band Z_{dr} becomes insensitive to
550 DSD because it is more dominated by small drops, which tend to be more spherical. In
551 particular, for DSDs with $D_m < 1$ mm, the CHILL S-band Z_{dr} was nearly 0 dB whereas Z_h
552 showed more noticeable variation with D_m . For DSDs with $D_m < 1$ mm Z_{dr} , which was around 0

553 dB (see Fig. 9a), does not seem to provide any useful information to retrieve the small end of the
554 drop size spectrum. Hence, a formula combining both parameters would be more appropriate for
555 D_m estimation. This would be particularly applicable for rain regimes which are dominated by
556 small drops, even at high rain rates, such as hurricane systems (e.g., Tokay et al. 2008, Brown et
557 al., 2016) as well as warm shallow rain in sub-tropical (e.g., Thurai and Bringi 2008) and
558 tropical oceanic locations (Thompson et al. 2015).

559

560 *c) GPM DPR retrieval algorithm*

561

562 The σ_M variation with D_m can be useful for the satellite-radar based estimates of rainfall rates at
563 ground level. Specifically, the GPM DPR needs to make assumptions on the shape of the DSD
564 in order to retrieve DSD parameters such as D_m and N_w (Iguchi et al. 2016). These findings,
565 along with our modified σ_M versus D_m variation, suggest that a variable (or more flexible) $\mu - D_m$
566 relationships be used in the satellite retrieval of the DSD parameters (if indeed gamma DSD is
567 assumed, rather than generalized gamma, as mentioned earlier). However, an initial assessment
568 of the DPR performance indicates the retrievals discussed above for the Huntsville event agree
569 within the limits of uncertainty. Preliminary comparisons between 2ADPR and GPM Ground
570 Validation Network (VN) DSDs, which rely on ground-based polarimetric radar data (e.g.,
571 WSR-88DP) to estimate D_m , suggest that the DPR and VN D_m 's associated with stratiform
572 precipitation are quite similar. The DPR estimates being biased only 0.1 mm high relative to VN
573 estimates. The mean absolute error of the DPR D_m retrievals relative to the ground radar
574 retrievals is 0.2 mm. Hence, the DPR retrieved D_m (1.8 mm) for this Huntsville event compares

575 rather well with the combined MPS-2DVD measurement of D_m (1.6 mm), at least within the
576 uncertainty of the 2ADPR D_m retrieval.

577

578 **6. Summary**

579

580 Two collocated disdrometers have been used to measure the full drop size distribution with high
581 resolution (50 μm) for small drops (MPS) and good resolution (170 μm) for moderate to large
582 drops (2DVD) in two spring-time rain events occurring in widely different climatologies. After
583 time integration of 1-2 h, the 2DVD-based spectra were found to systematically underestimate
584 the concentrations at the small drop end relative to the MPS-based spectra. There was very good
585 agreement in the overlapping size interval between the two instruments giving confidence in the
586 interpretation of the combined spectra in terms of physical processes as opposed to instrument-
587 to-instrument differences. Examination of the combined spectra revealed a drizzle mode for
588 $D < \sim 0.7$ mm and a precipitation mode for larger diameters in agreement with the identification of
589 such modes by Able and Boutle (2012) which was based on using aircraft imaging probes (2D-C
590 and 2D-P) in warm rain oceanic clouds. While the two events analyzed herein were from
591 different regions (Greeley, CO and Huntsville, AL), the two modes could be easily identified in
592 the combined spectra (largely independent of rain rate). However, no attempt is made here to
593 suggest physical processes giving rise to the two modes other than the general domination of ice-
594 phase microphysics in the Greeley event and warm cloud base convection with component of
595 warm rain microphysics in the Huntsville event, with negligible evidence of evaporation causing
596 a depletion of tiny drops at either location as inferred from the presence of the drizzle mode
597 throughout the duration of the precipitation events.

598

599 The 1-min averaged combined spectra were also analyzed in terms of the parameters N_w , D_m and
600 σ_M which are relevant for radar applications (based on both polarimetric and dual-frequency).
601 Note that these three parameters are defined in terms of higher order moments of the spectra (3rd
602 moment and/or higher moments) with no assumption of the gamma DSD model (Haddad et al.,
603 1996, Ulbrich and Atlas, 1998, Testud et al. 2001). While all three parameters are affected by the
604 small drop concentrations especially at light rain rates, the N_w and the ratio σ_M/D_m were found to
605 be significantly affected (significantly larger) by the small drop concentrations in the drizzle
606 mode even at high rain rates in the two events analyzed herein. This result is particularly relevant
607 for radar-based retrievals which assume the gamma model (the parameters being N_w , D_m and the
608 shape μ) with the μ parameter being fixed (≈ 3) or based on D_m . The general tendency (under
609 such assumptions) is for the radar retrievals to overestimate D_m and underestimate N_w . Clearly,
610 more data with the combined MPS-2DVD instruments are needed in a variety of rain rates and
611 different climatologies to improve the radar-based retrievals. Such datasets should also impact
612 the numerical modeling of the microphysics of rain processes which use multi-moment bulk
613 schemes (here the total number concentration is of primary importance and it is obvious that the
614 drizzle mode in the combined DSD would play a significant role).

615

616 For each of the two events analyzed, we also had available polarimetric radar data and were able
617 to characterize the variation of radar measured Z_{dr} (and Z_h) with D_m from the combined spectra.
618 In the Greeley event, this variation was well-captured at the small values of D_m (< 0.5 mm)
619 where measured Z_{dr} tended to 0 dB which precluded the estimation of D_m based on Z_{dr} alone. A

620 retrieval of D_m using both Z_h and Z_{dr} would be more appropriate but will be addressed as more
621 combined datasets become available in the future.

622
623 The small raindrop findings presented here also have implications for satellite-based retrieval of
624 DSD parameters and ultimately surface rainfall rates. An overpass of the GPM DPR during an
625 event of the Huntsville campaign revealed that the DPR algorithm overestimated D_m and
626 underestimated N_w relative to the combined MPS-2DVD measurements. This indicates that a
627 fixed μ in the gamma distribution may not be the most fitting assumption to describe the DSD
628 from a satellite-based radar perspective. Instead, the above σ_M - D_m relationship, which for
629 application purposes needs to be converted into μ - D_m space (e.g., Williams et al. 2014), could
630 facilitate more accurate retrievals. However, this speculation requires further investigation since
631 unlike Williams et al. (2014), we did not account for any correlation that might exist between σ_M
632 and D_m calculated from the combined MPS-2DVD measurements before fitting a power-law.

633
634 Although only two events are reported in this paper, analyses of several other events have also
635 shown that the full DSD spectra has the aforementioned drizzle mode and the precipitation mode
636 which together could be better represented by the use of generalized gamma function (Auf Der
637 Maur, 2001, Lee et al. 2004).

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643 swath across northern Alabama during the 11 April 2016 event, from the 2ADPR product.

Appendix

Calculation of DSDs from the 2DVD data is a very well established procedure and hence only the calculation of DSDs from the MPS data is summarized here. This involves the sample area calculation which is described in the MPS data analysis guide by Droplet Measurement Technologies. It first entails the calculation of the effective array width (EAW) and the depth of field (DoF).

$$EAW = R_p * (n - x - 1) \quad (A1)$$

where R_p is the probe resolution, n is number of diodes (= 64) and x is the bin number (1:62), and

$$DoF = 6 r^2 / \lambda \quad (A2)$$

where r is the particle radius and λ is the laser wavelength, all in MKS units.

The DSD denoted by $N(D)$ is then given by:

$$N(D) = C / (A_{eff} * v * \Delta t * \Delta d) \quad (A3)$$

where C is the number of particles measured in the diameter interval Δd , and Δt is the time interval and v represents the particle velocity, and A_{eff} is the effective area given by:

$$A_{eff} = EAW * DoF \quad (A4)$$

Note although the MPS measures the velocity of each particle, in our DSD calculations, we have used a recommended velocity-diameter relationship for small drops:

$$v = -19.27 + (0.50 * D_\mu) - (9.04 * 1e-5 * D_\mu^2) + (5.66 * 1e-9 * D_\mu^3) \quad (A5)$$

which yields v in cm/s at sea level and requires a suitable correction factor for high altitudes such as Greeley, and D_μ is D_{eq} in microns. Eq. (A5) yields similar values to the Gunn-Kinzer (1949) data for drops smaller than 1.2 mm.

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858

859 FIGURE CAPTIONS

860
861 FIG. 1: (a) The MPS, 2DVD and Pluvio inside the double wind fence at the site near Greeley,
862 Colorado (40.3273569N, 104.6093944W, 1.4 km AMSL). (b) MPS, 2DVD and Parsivel
863 disdrometers inside the double wind fence at the Huntsville site (34.7233333N, 86.6419444W,
864 212 m AMSL).

865
866 FIG. 2: (a): Picture of a Meteorological Particle Spectrometer and the custom-design stand. A
867 wind vane aligns the sample path with the wind flow but was not used in our campaigns since the
868 instrument was installed within a DFIR. (b): schematic of a drop falling through the MPS sensor
869 measurement area; There are 64 photo-detectors and the horizontal resolution is 50 μm . From
870 Droplet Measurement Technologies (DMT).

871
872 FIG. 3: (a) Fall velocity versus drop equi-volume diameter (D_{eq}) from the 2DVD data as 2D
873 frequency of occurrence plot. The dashed line represents the equation given in Atlas et al. (1973)
874 that approximates the Gunn-Kinzer terminal fall speed measurements (Gunn and Kinzer 1949),
875 and the dotted line is this approximation after applying altitude correction for the 1.4 km AMSL
876 for Greeley. (b) Velocity of histograms specific to all drops with D_{eq} values of 2.5 ± 0.1 mm. The
877 expected values at sea level (7.3 m/s) and at 1.4 km altitude (7.9 m/s) are shown as dashed line
878 and dot-dash line respectively.

879
880 FIG. 4: (a) 1-minute rain-rate (R) from Pluvio for the entire event; (b) the corresponding rain
881 accumulation; (c) the corresponding 2-hour rain accumulations.

882

883 FIG. 5: 2-hour DSD comparisons from the 2DVD (crosses) and 2DVD-MPS combined
884 (diamonds) for the 17 April 2015 event. The time interval is specified for each case. Note log-log
885 scale is used to focus on the small drops.

886

887 FIG. 6: (a) D_m derived from 1 minute DSDs using 2DVD data (grey points), and the combined
888 MPS-2DVD data (black points); (b) the corresponding σ_M values; (c) the σ_M - D_m scatter plot
889 using the same DSDs and their fitted curves for $D_m \geq 0.5$ mm. 3-minute smoothing is applied to
890 (a) and (b) in order to show more clearly the differences in grey and the black points.

891

892 FIG. 7: D_m histogram comparisons for three different rain types (as indicated in the top panels)
893 and for different rainfall rate intervals (as indicated in bottom panels). The 2DVD data-based
894 histograms are shown in grey and the combined MPS-2DVD DSD based histograms are shown
895 in black. All histograms are based on 1-minute DSDs.

896

897 FIG. 8: D_m comparisons between 2DVD-based and 2DVD-MPS combined DSD based estimates.
898 Each data point is based on 1-min spectra.

899

900 FIG. 9: (a) the CHILL S-band Z_{dr} and (b) the CHILL S-band Z_h measurements over the
901 instrument site versus D_m calculated using 2DVD (grey) and MPS-2DVD combined DSDs
902 (black).

903

904 FIG. 10: (a) Fall velocity versus drop D_{eq} from the 2DVD data as 2D frequency of occurrence
905 plot. The dashed line represents the equation given in Atlas et al. (1973) which represents the
906 Gunn-Kinzer variation, (b) histogram of vertical velocity specific to all drops with D_{eq} values of
907 2.5 ± 0.1 mm. The expected value at sea level (7.3 m/s) is shown as dot-dash line.

908

909 FIG. 11: Hourly DSD comparisons from the 2DVD (crosses) and 2DVD-MPS combined
910 (diamonds) for the 11 April 2016 event in Huntsville. The time interval is specified for each
911 case. The hourly rain accumulations were 1.3, 2.3, 1.6, 4.1, 4.4, and 3.7 mm for the 17, 18, 19,
912 20, 21, 22 hr UTC respectively.

913

914 FIG. 12: (a) D_m derived from 1 minute DSDs (after smoothing over 3 minutes) using 2DVD data
915 alone (grey circles), and the combined MPS-2DVD data (black crosses); (b) the corresponding
916 σ_M values; (c) the corresponding $\log_{10}(N_w)$.

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919 corresponding Z_{dr} (black crosses) and values of D_m derived from the combined DSDs; (c)
920 variation of the C-band Z_{dr} with D_m values from the 2DVD DSDs (grey) and the combined DSDs
921 (black). Note that some of the scatter is due to radar measurement error.

922

923 FIG. 14: The GPM DPR swath across northern Alabama during the 11 April, 2016 event
924 showing a) D_m [mm] and b) $10 \cdot \log_{10}(N_w)$ [$m^{-3}mm^{-1}$] both at 500-m AGL from the 2ADPR
925 product, and (c) the 10-minute DSD from 23:25 – 23:35 UTC, from the MPS and 2DVD around
926 the GPM overpass time period, at 23:31 UTC..

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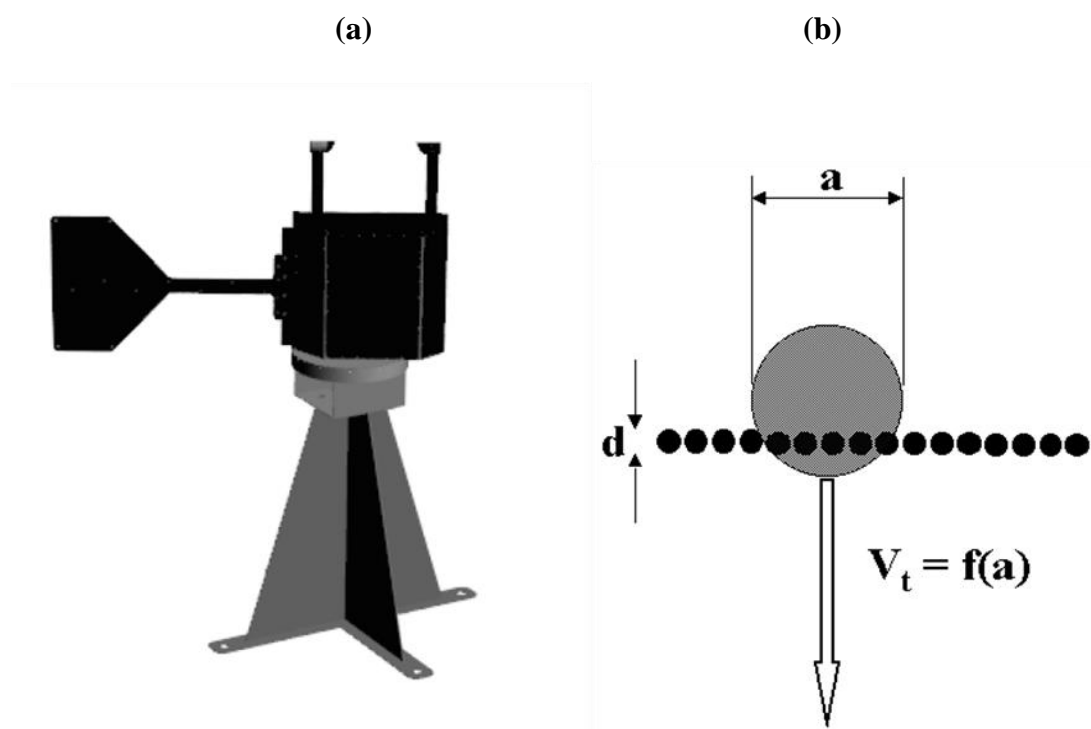
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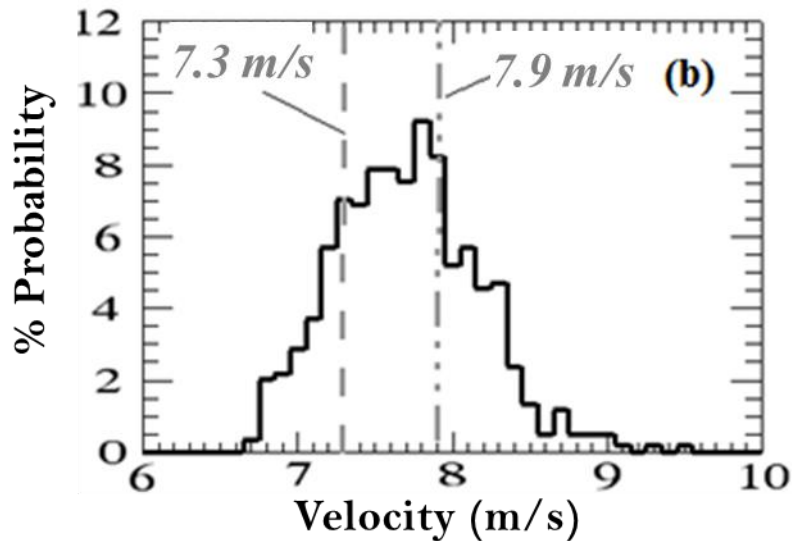
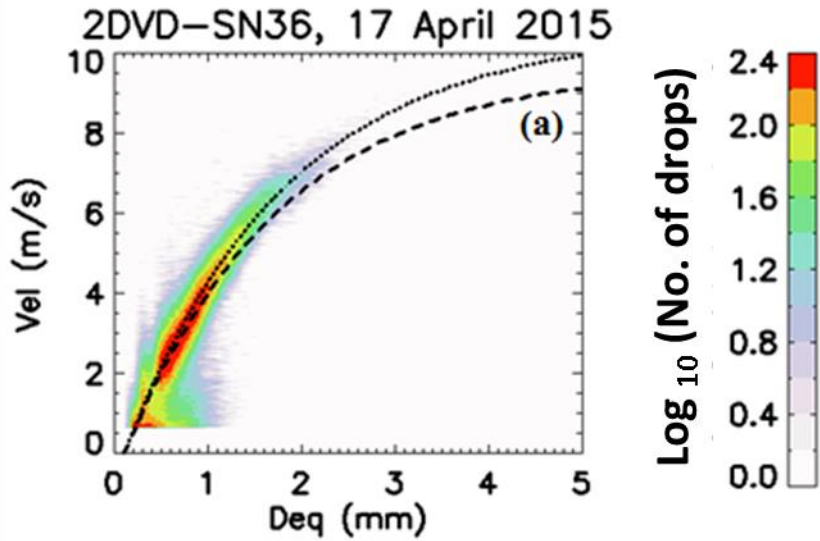
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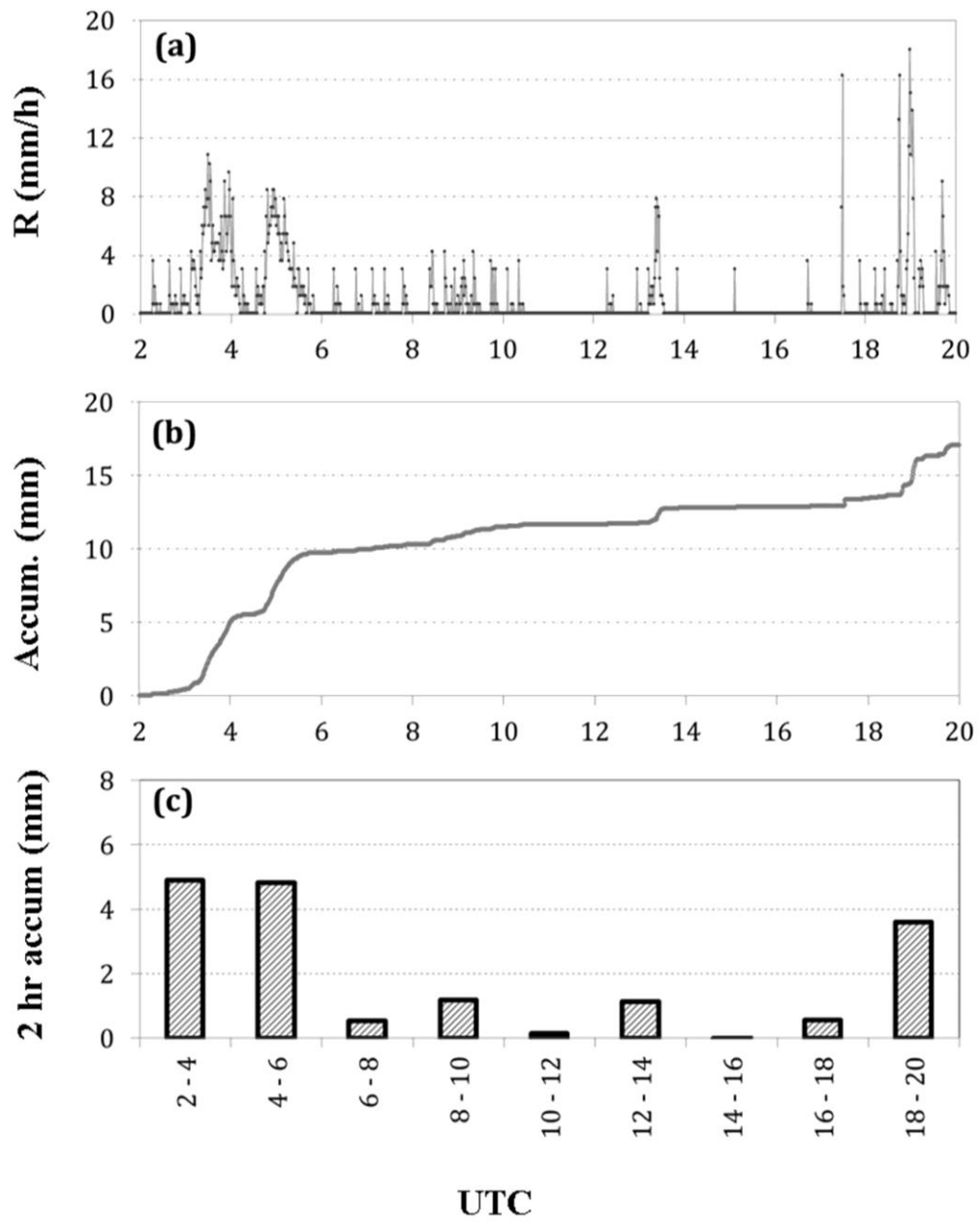
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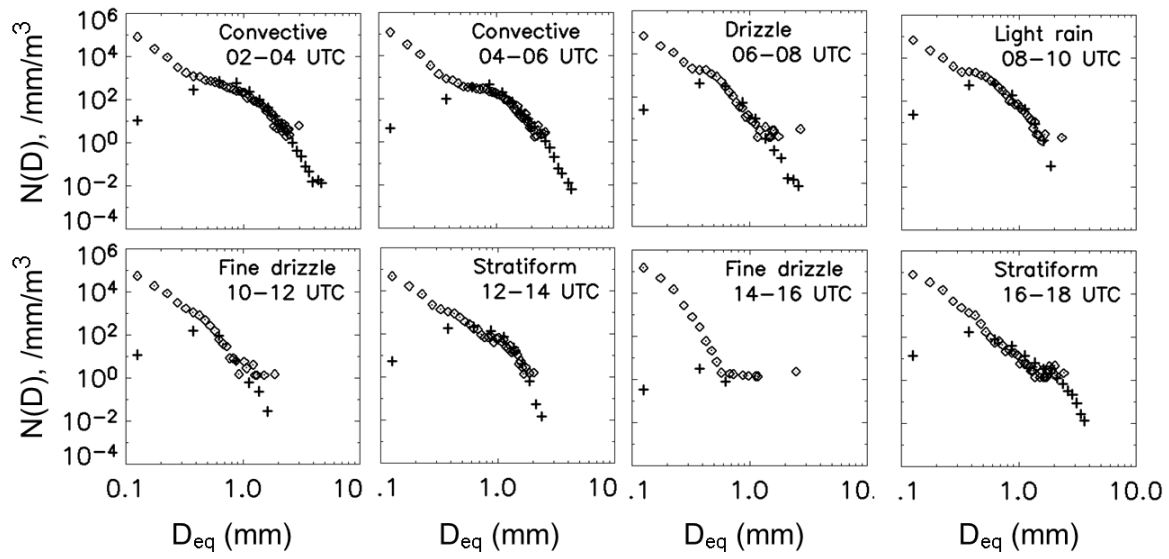


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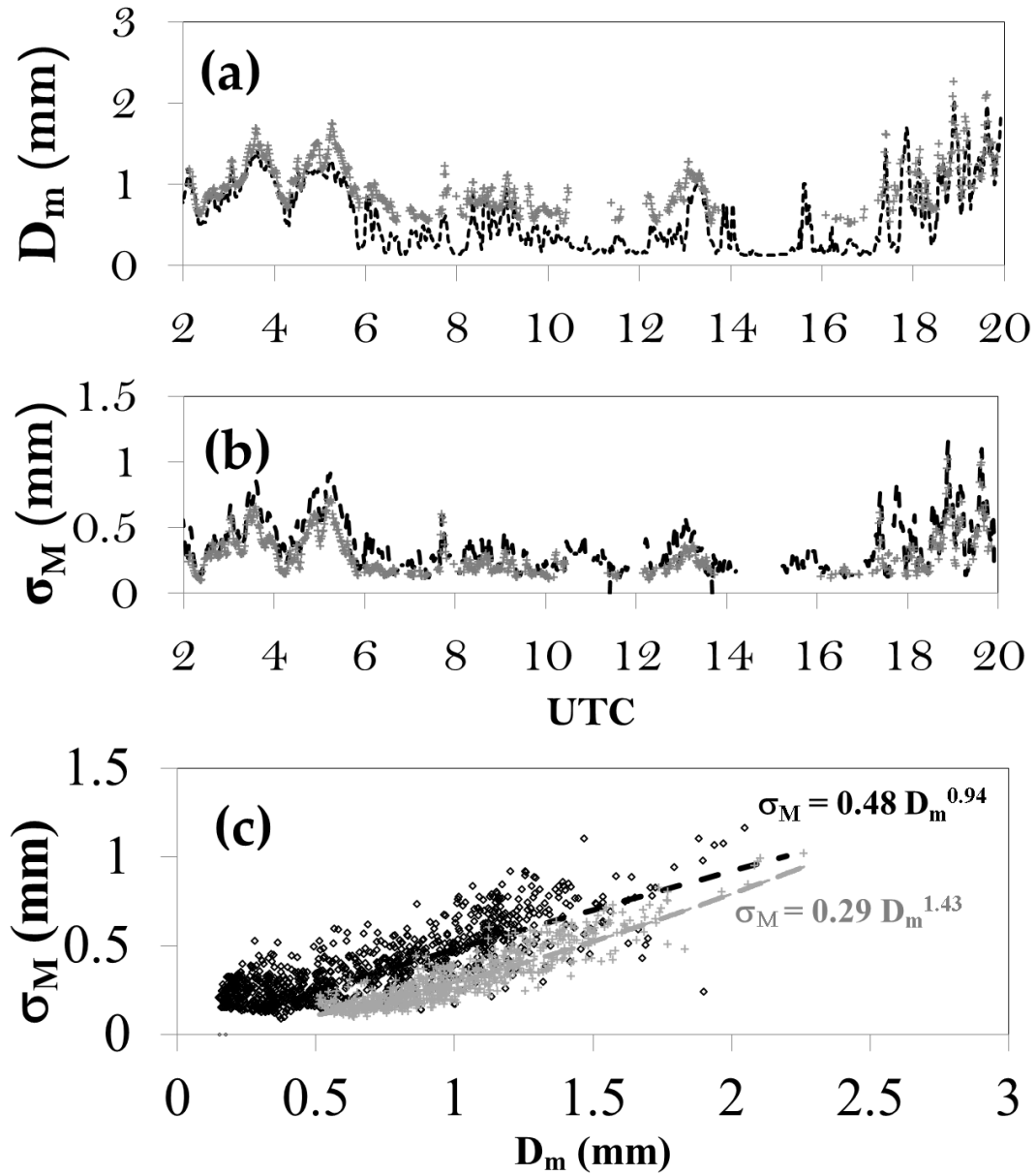
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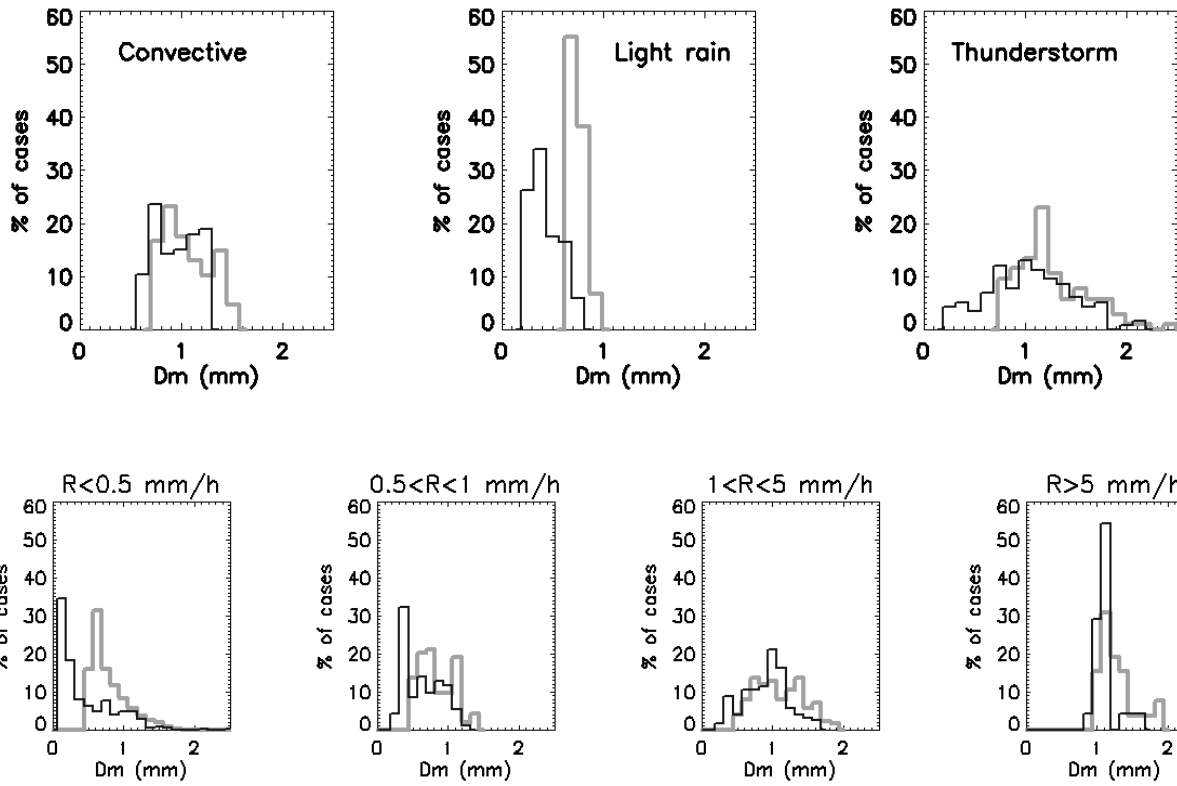
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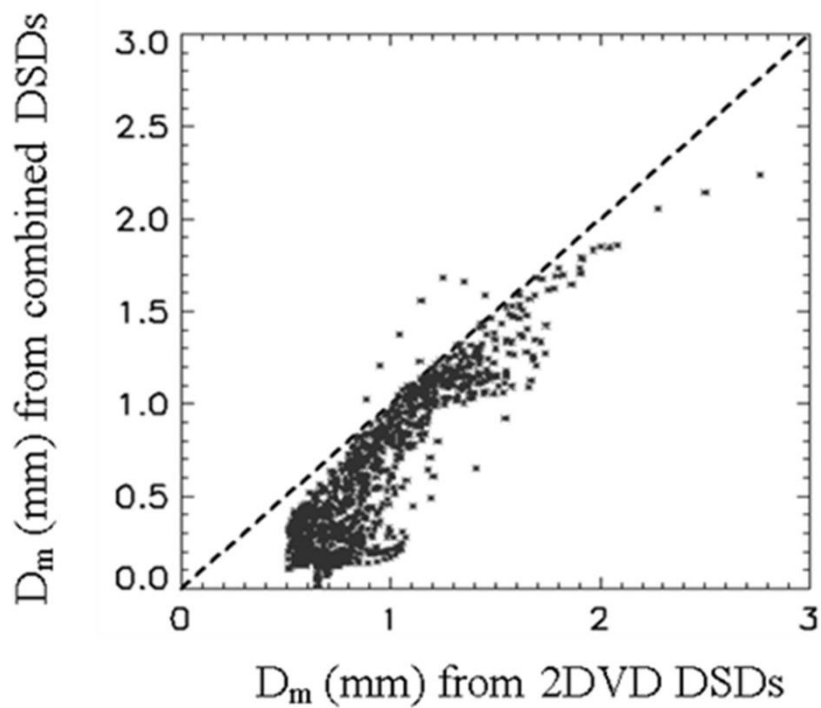
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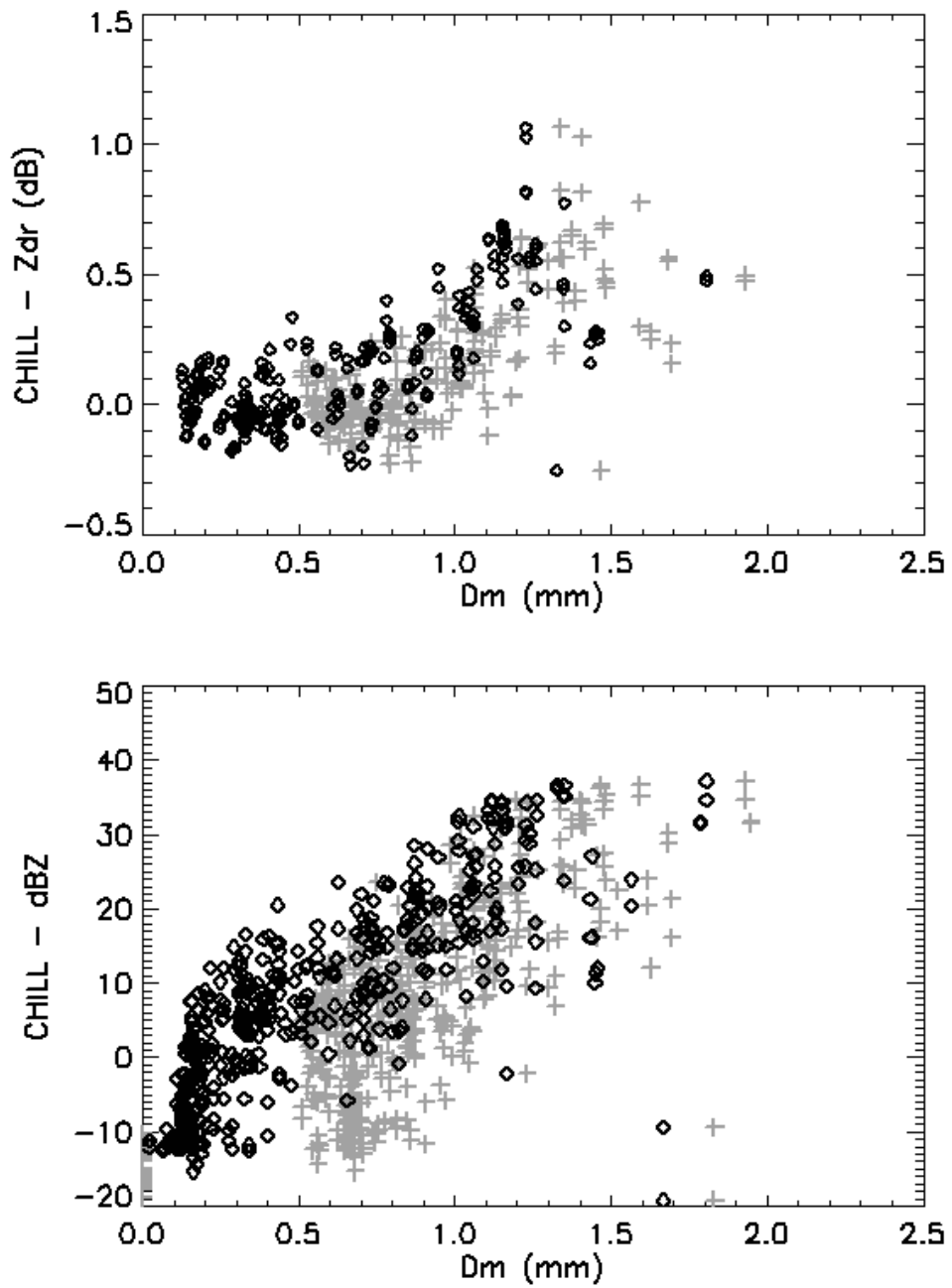
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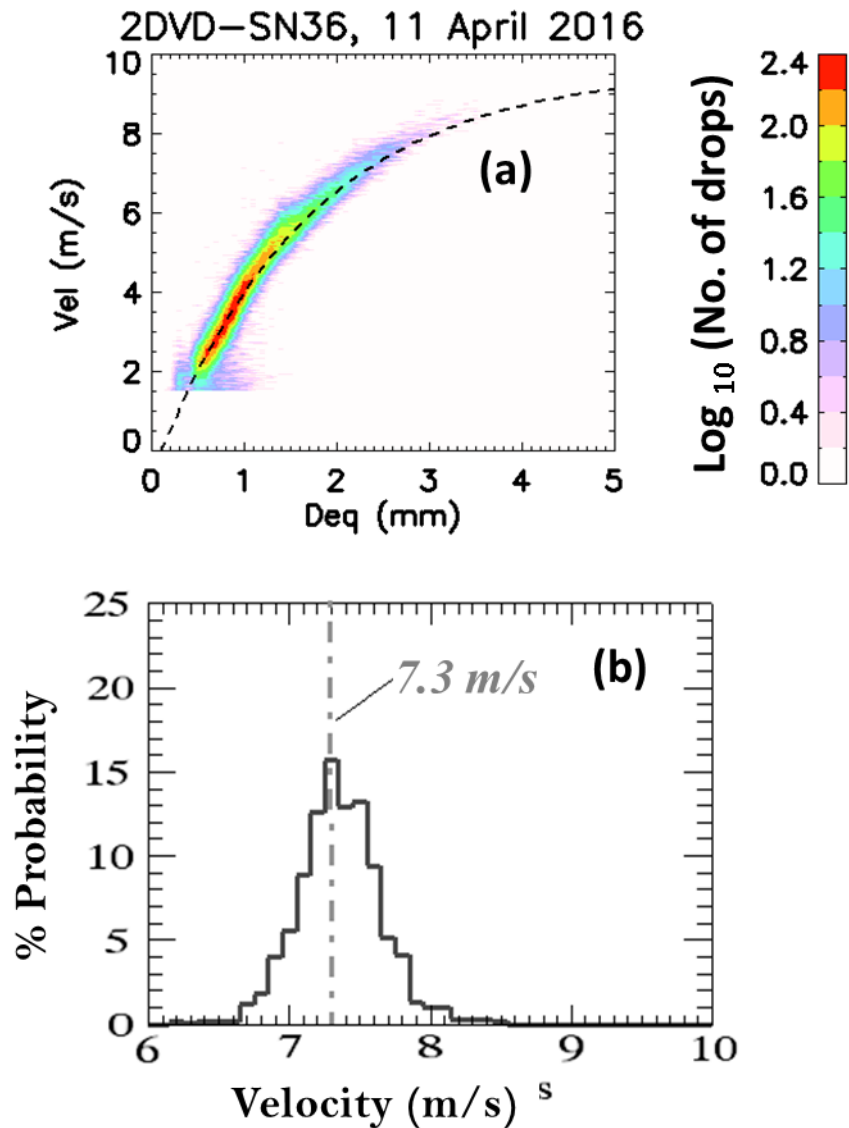


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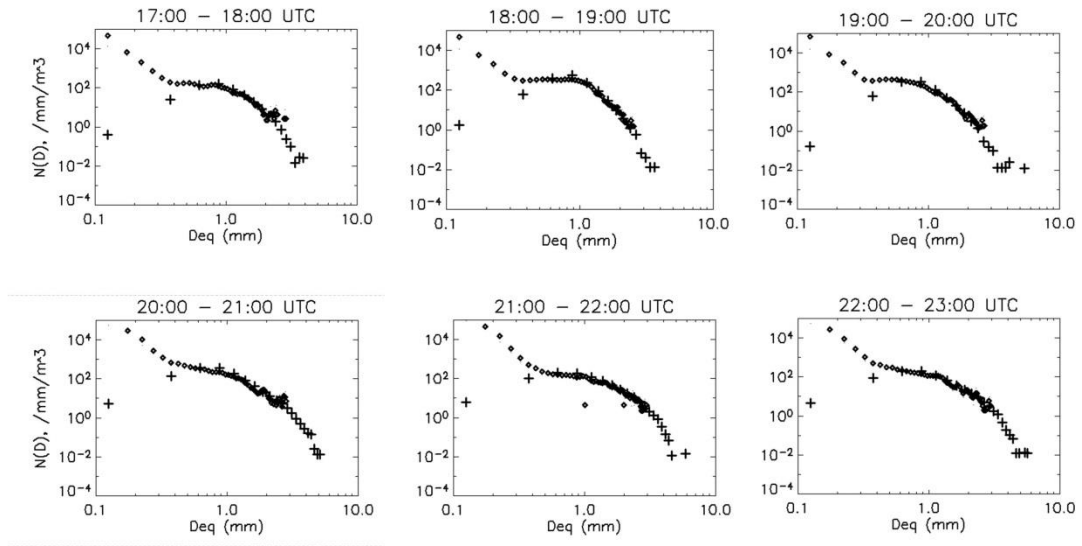
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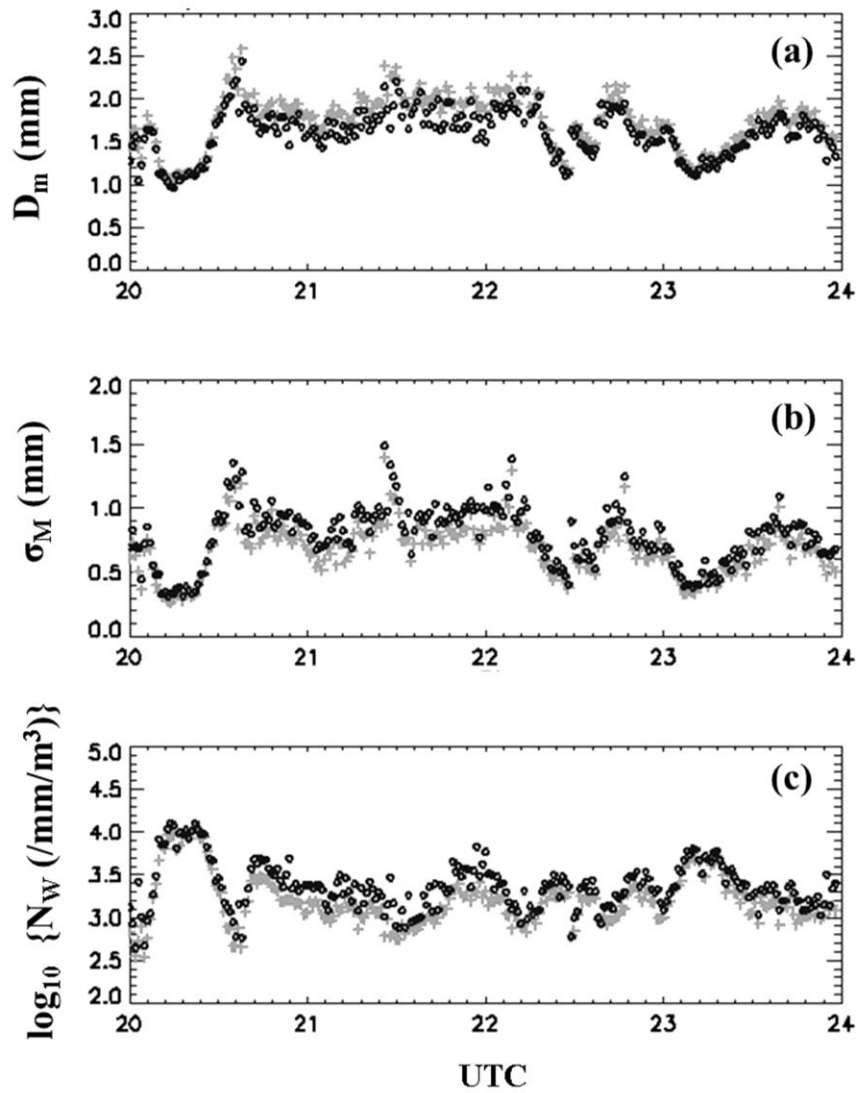
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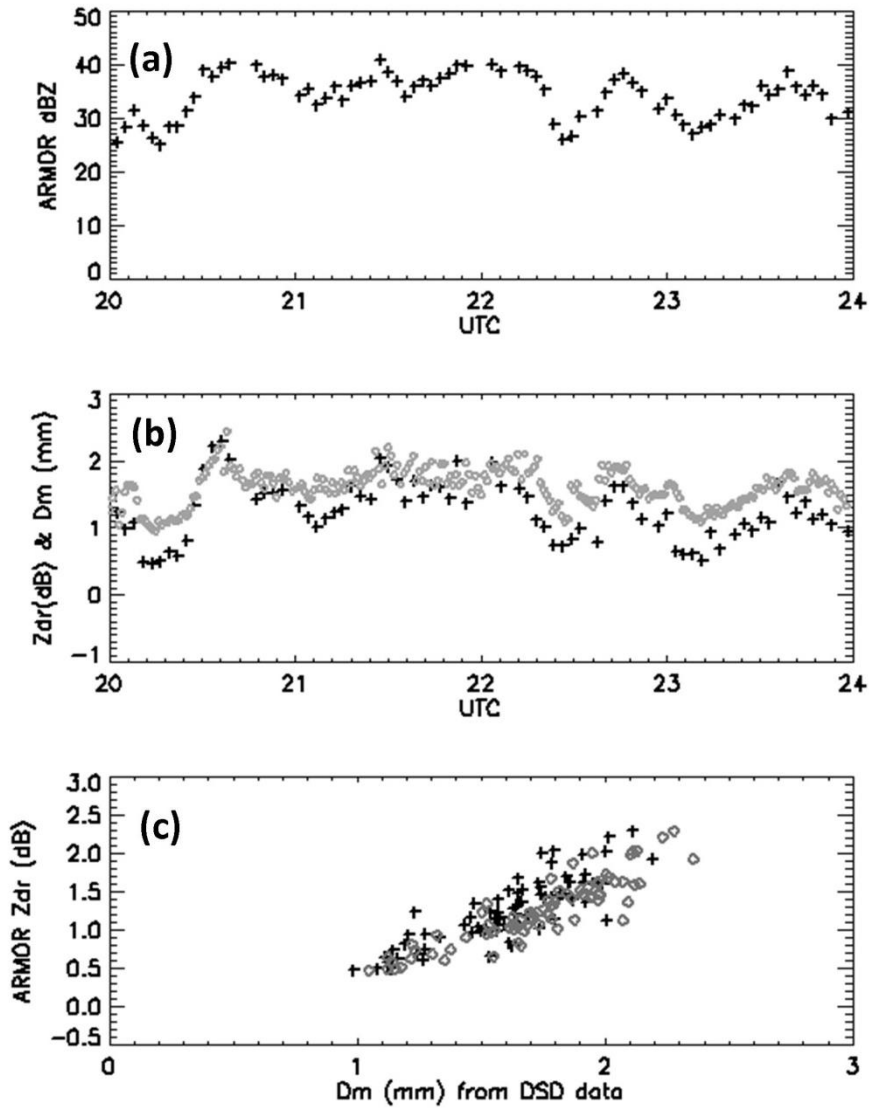
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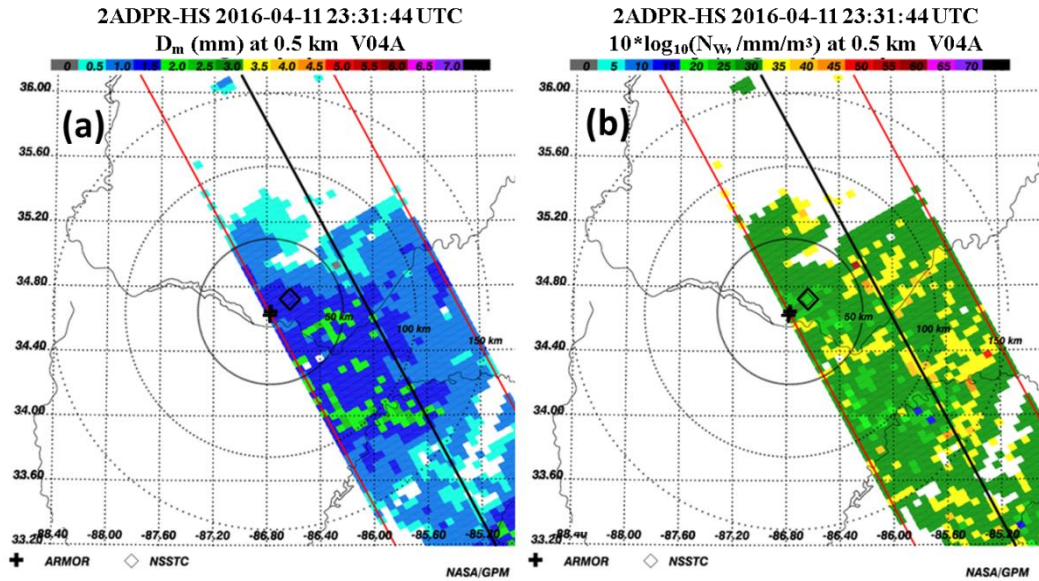
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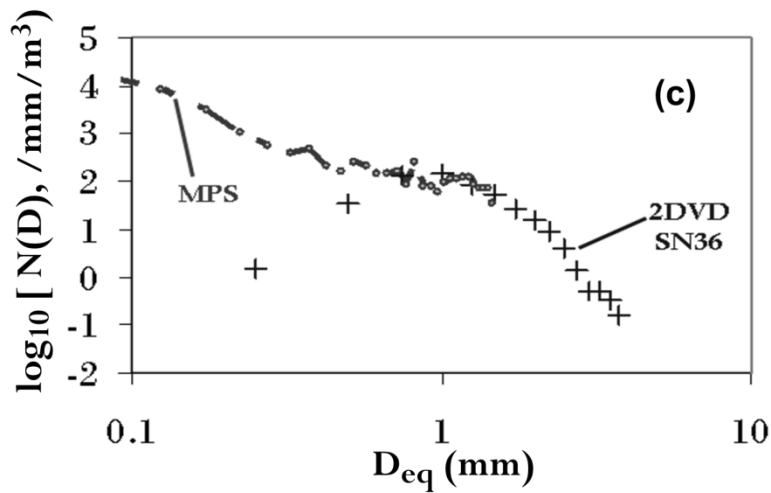
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1035 product, and (c) the 10-minute DSD from 23:25 – 23:35 UTC, from the MPS and 2DVD around

1036 the GPM overpass time period, at 23:31 UTC.

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TABLE 1: Specifications of the MPS and 3rd generation low-profile 2DVD (SN36)

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Parameter	MPS	2DVD (SN36)
Number of active pixels	62	625–630
Clock/line scan frequency	Selectable, max 200 kHz (when matched to fall speed of 10 ms ⁻¹)	55.172 kHz
Horizontal resolution	50 μm	170 μm
Measuring area	20 cm × 3.1 cm	10 cm × 10 cm
Vertical resolution	50 μm	Depends on fall speed (100 μm for 5 ms ⁻¹)
Size range	50 μm–3.1 mm	Typically > 0.6 mm
Calibration	Spinning glass disk with opaque dots of known size	Distance between vertical planes (i.e., plane test). Metal calibration spheres of known size

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1044 TABLE 2: Dominant rain types during the two hour periods for the 17 April 2015 event,

1045 classified using CHILL RHI scans over the disdrometers.

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2 hour period	Dominant Rain type
00 – 02 UTC	Dominated by melting snow (not included in the DSD analyses)
02 – 04 UTC	Moderately strong convective rain
04 – 06 UTC	Moderately strong convective rain
06 -08 UTC	Drizzle
08 – 10 UTC	Light rain
10 – 12 UTC	Fine drizzle
12 – 14 UTC	Mostly stratiform rain
14 – 16 UTC	Fine drizzle
16 – 18 UTC	Mostly stratiform rain
18 – 20 UTC	Thunderstorm

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1050 TABLE 3: Two-hour rain accumulations for the three convective rain periods in Table 2

2-hr	SN36 (mm)	SN36 & MPS (m)	Pluvio (mm)
2-4	4.85	4.94	4.91
4-6	4.11	4.51	4.84
18-20	2.79	2.94	3.62

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